

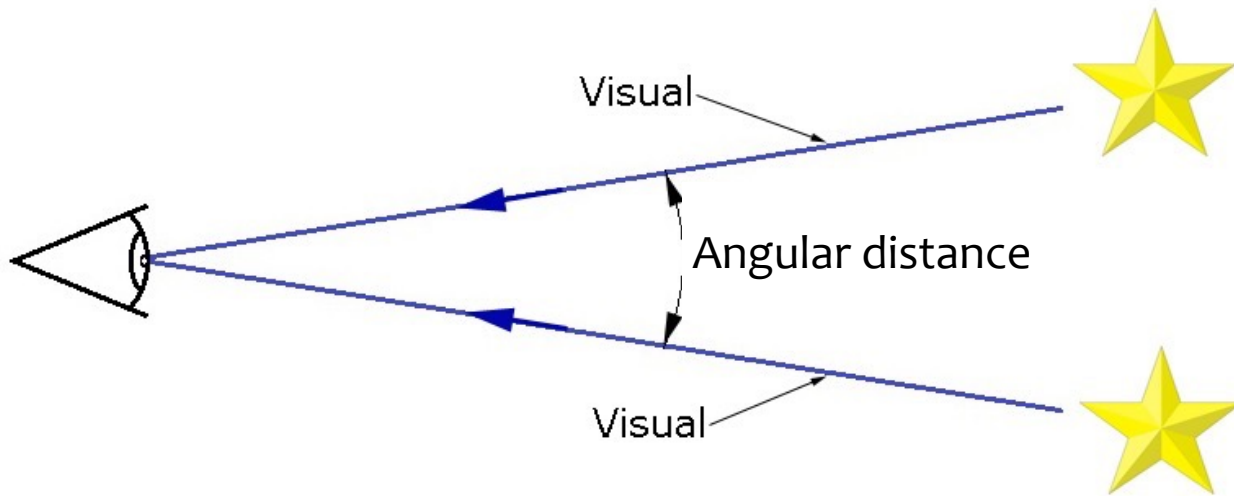
University of Texas Rio Grande Valley
Center for Gravitational Wave Astronomy
Nompuewenu Observatory

Lesson 5

Brownsville - 2016

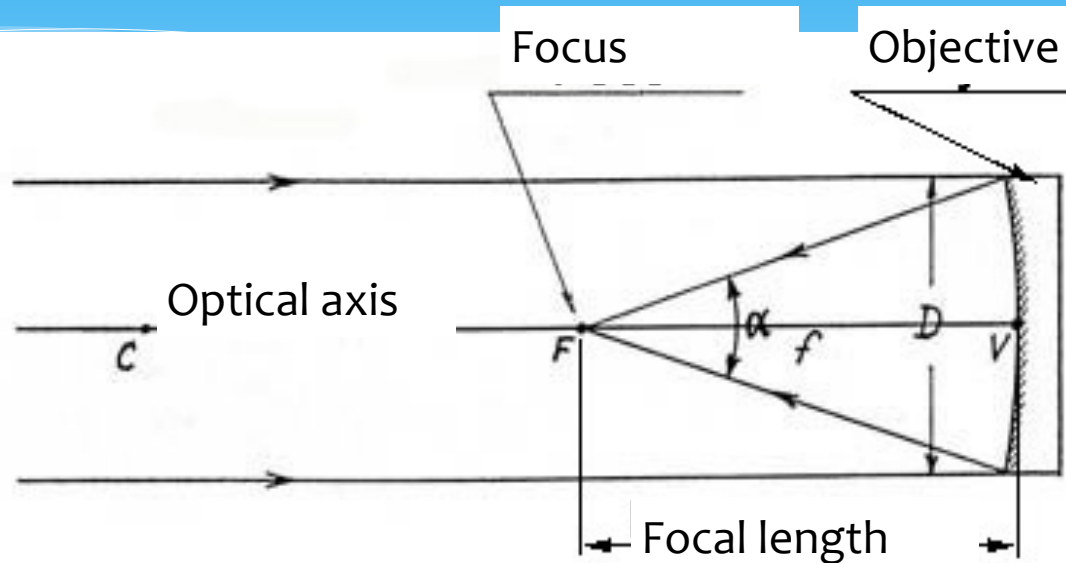
Visual

- * Visual.
- * Angular distance
- * Angular diameter
- * Examples

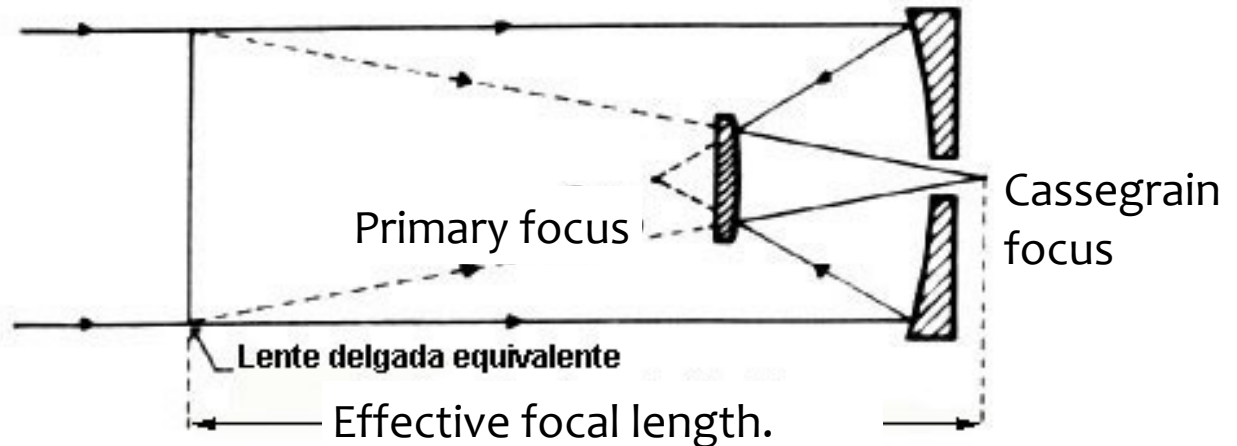


Telescope Parameters

- * Objective.
- * Focus.
- * Focal plane.
- * Focal length.
- * Effective focal length.



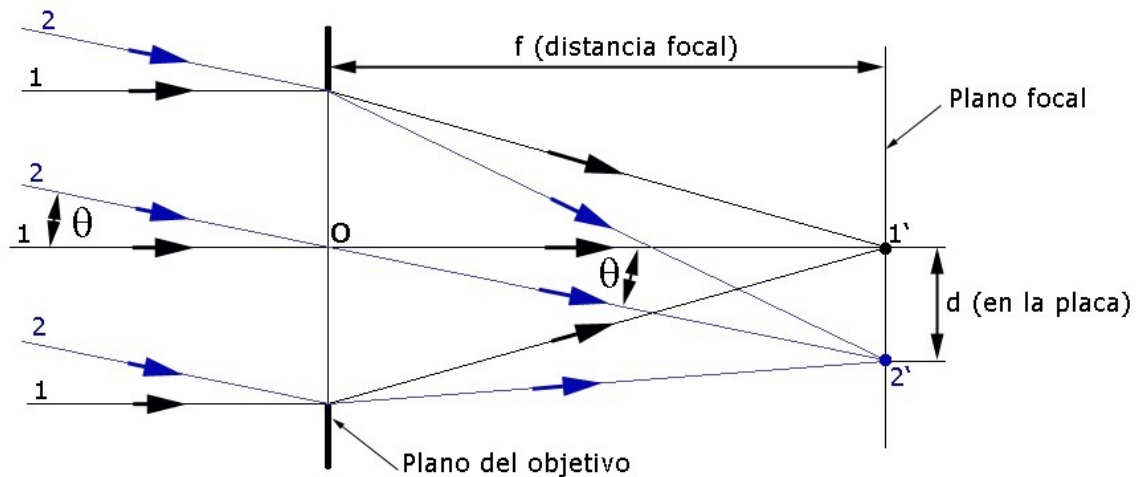
- * Example:
 - MEADE 16" LX200
 - $f = 4064 \text{ mm}$



CCD parameters

- * Pixel size
- * Sensor size
- * Example:
 - APOGEE F16M:
 - Pixels: Width = 9 mm y height = 9 mm.
 - Sensor: Width = 4096 píxels Height = 4096 píxels.
 - Sensor dimensions:
 - Width = 36,8 mm Height = 36,8 mm.

Plate field



Field Of View – FOV.

$$\theta = d / f \text{ [radianes]}$$

$$\theta = 180/\text{PI} * d / f \text{ [}^\circ\text{]}$$

$$\theta = 206264,8 * d / f \text{ [arcseg]}$$

Example

* APOGEE F16M in MEADE 16" LX200

$$\theta(\text{width}) = 206264,8 * 36,86 \text{ mm} / 4064 \text{ mm}$$

$$\theta(\text{width}) = 1870''$$

$$\theta(\text{width}) = (1870/60)' = 31,2'$$

$$\theta(\text{height}) = 206264,8 * 36,86 / 4064 \text{ mm}$$

$$\theta(\text{height}) = 1870''$$

$$\theta(\text{height}) = (1870/60)' = 31,2'$$

Plate scale

- * Plate scale in chemical photography:

$$\text{Plate scale} = \theta / \text{unit of length. [arseg/mm]}$$

- * Plate scale in CCDs:

$$\theta = 206264,8 d / f, \text{ d is the pixel dimensión.}$$

$$\text{Pixel scale} = 206264,8 * d / f \text{ [arcsec/pixel]}$$

- * Example: MEADE 16'' with APOGEE F16M:

$$\text{Plate scale} = 206264,8 * 0,009 \text{ mm} / 4064 \text{ mm}$$

$$\text{Plate scale} = 0,46 \text{ ''/pix}$$

Space resolution

- * Space resolution = # Píxels / angular unit
- * Space resolution = 1 / Plate scale [pixels/arcsec]
- * Example MEADE 16'' with APOGEE F16M:

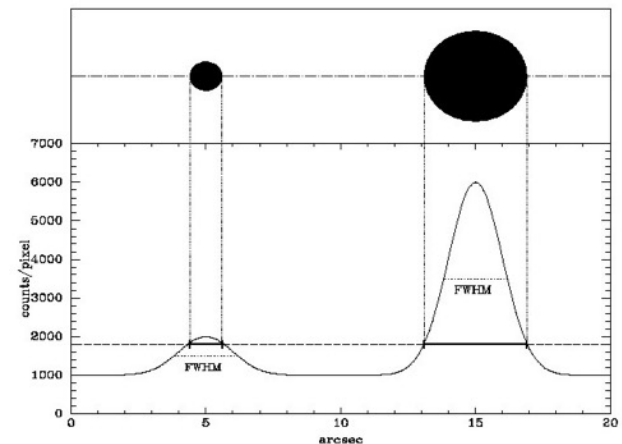
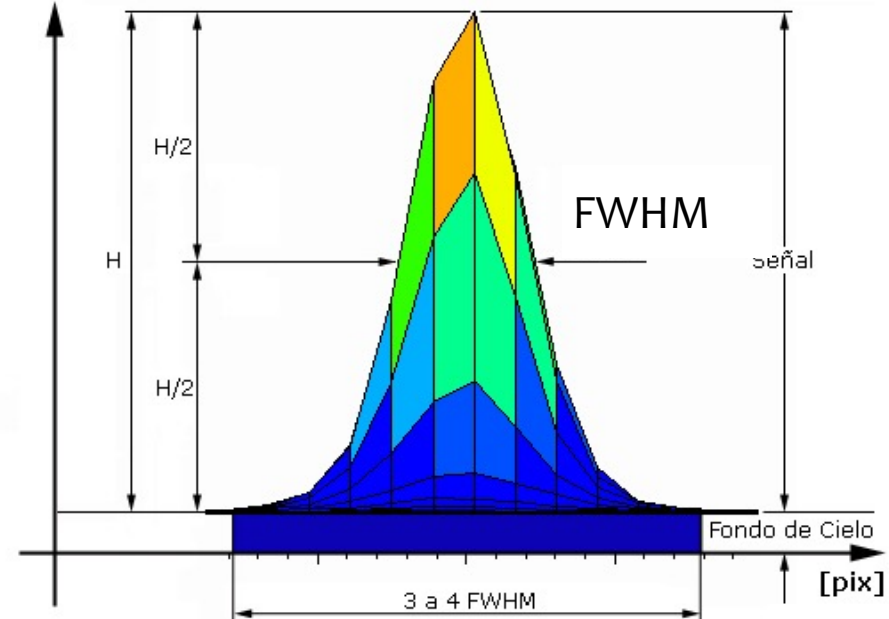
$$\text{Space resolution} = 1 / (0,46 \text{ ''/pix})$$

$$\text{Space resolution} = 2,17 \text{ pixeles/arcsec}$$

Seeing

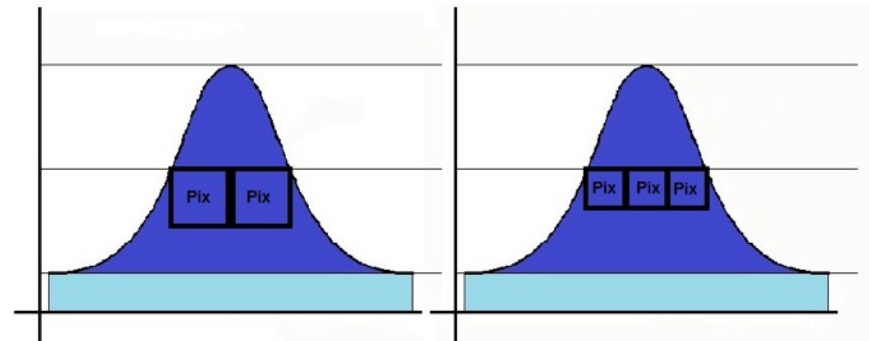
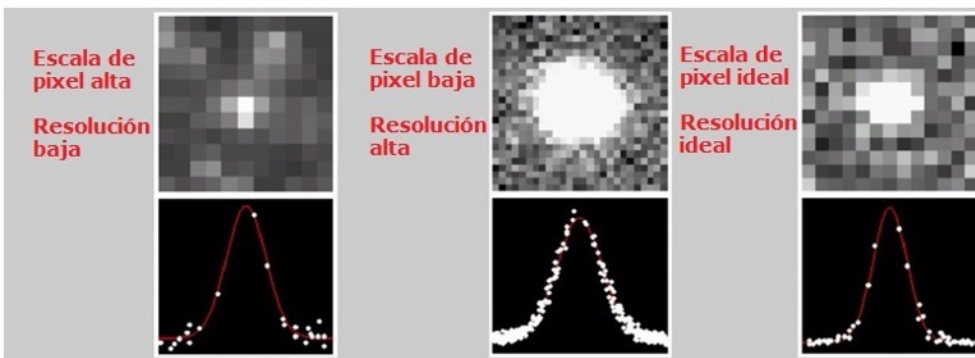
- * Atmospheric distortion.
- * Other effects: terrain, dome, instrument, etc.
- * PSF (Point Spread Function): how a point is spread on the detector
- * Seeing measurement: with the FWHM (Full width at Half Maximum).

Counts



Relationship between seeing and plate scale

- * Typical Seeing .
- * ideal plate scale .
- * Example: If a given seeing es FWHM = 5 arcsec
Pixel scale max= $5 \text{ arcsec} / 2 = 2,50 \text{ arcsec/pix}$
Pixel scale mín = $5 \text{ arcsec} / 3 = 1,66 \text{ arcseg/pix}$



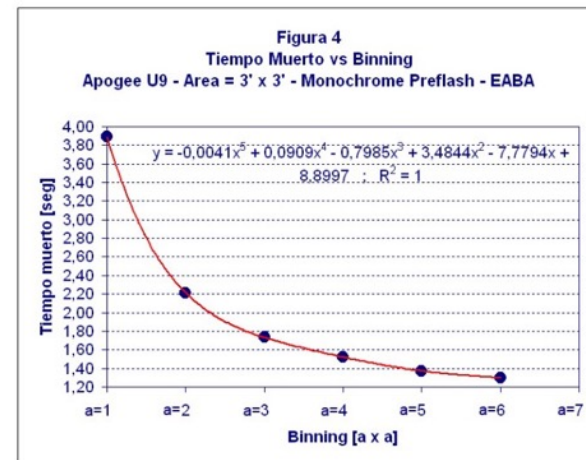
Binning and seeing

- * Binning: superpíxel.
- * Binning is done to find ideal plate scale
- * Example MEADE 16" with APOGEE F16M :

Binning 5x5 : Plate scale = 2,28 arcsec/pix

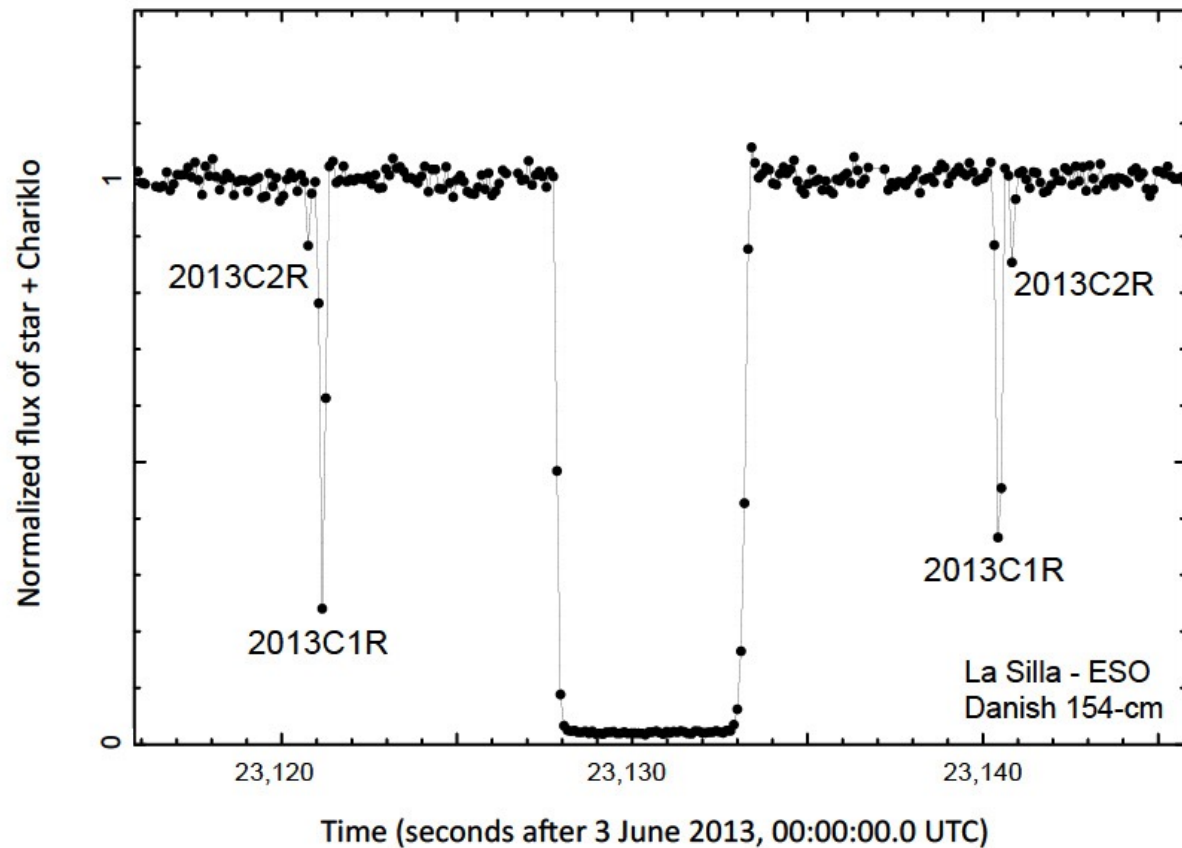
Binning 4x4: Plate scale = 1,83 arcsec/pix

Binning Options	Combined pixels on the CCD Chip
None	
2 x 2 (4 pixels = 1)	
3 x 3 (9 pixels = 1)	
4 x 4 (16 pixels = 1)	



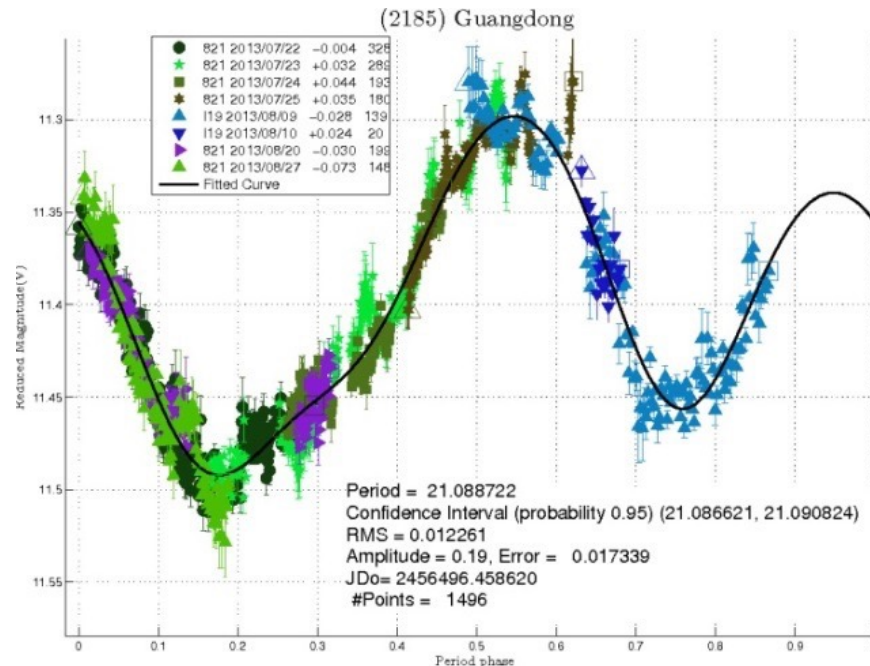
Binning and time resolution

* Time resolution = # images / time



Appropriate time exposure

- * Try best SNR (signal to noise ratio).
- * Don't go over linearity. Consider variation of zenith distance during observation.
- * Don't lose time resolution, when observing short duration events.

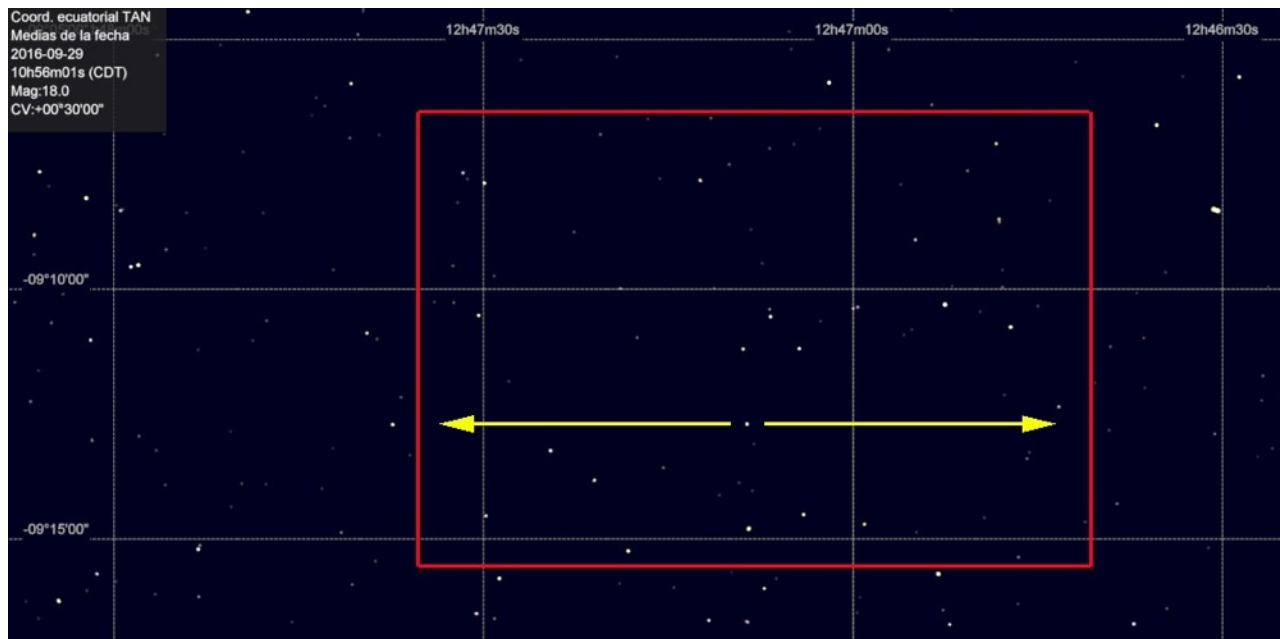


Telescope balancing (German equatorial mount)

1. Polar axis of mount: parallel to world axis.
2. Optical axis parallel to the meridian and braking on both axis.
3. Displacements: weights in the declination axis declinación and tube encasing.
4. Tube balancing: freeing declination brake and moving the tube
5. Polar axis: freeing polar axis brake and moving weights of declination axis.

Camera orientation

- * When moving scope in the E-W direction stars' traces must be parallel to the edge of the larger border of the rectangular FOV.



Cardinal points in the visible field

- * East of image: when moving scope to the east, East is the side of the image (up or down) where stars first appear.
- * North: North is the the side of the image where stars first appear when moving the scope North.

Stationing of equatorial mount

- * It is the fine adjustment of the mount polar axis position. The drift method will be applied in the declination.
- * Azimut adjustment: Look for a star close to the equator and about to culminate. If the star drifts to the North mount's North is to the West of of the actual North.
- * Altitude adjustment: Pick a star with right ascension -6h and declination approx 20° . If the star drifts to the North, the mount's north is higher than the celestial north.